

# **Dark Matter, Dark Energy & SNe with MOSAIC and Wider Field Cameras**

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# Science Drivers

- **DARK MATTER**: SuperMacho
- **DARK ENERGY**: ESSENCE

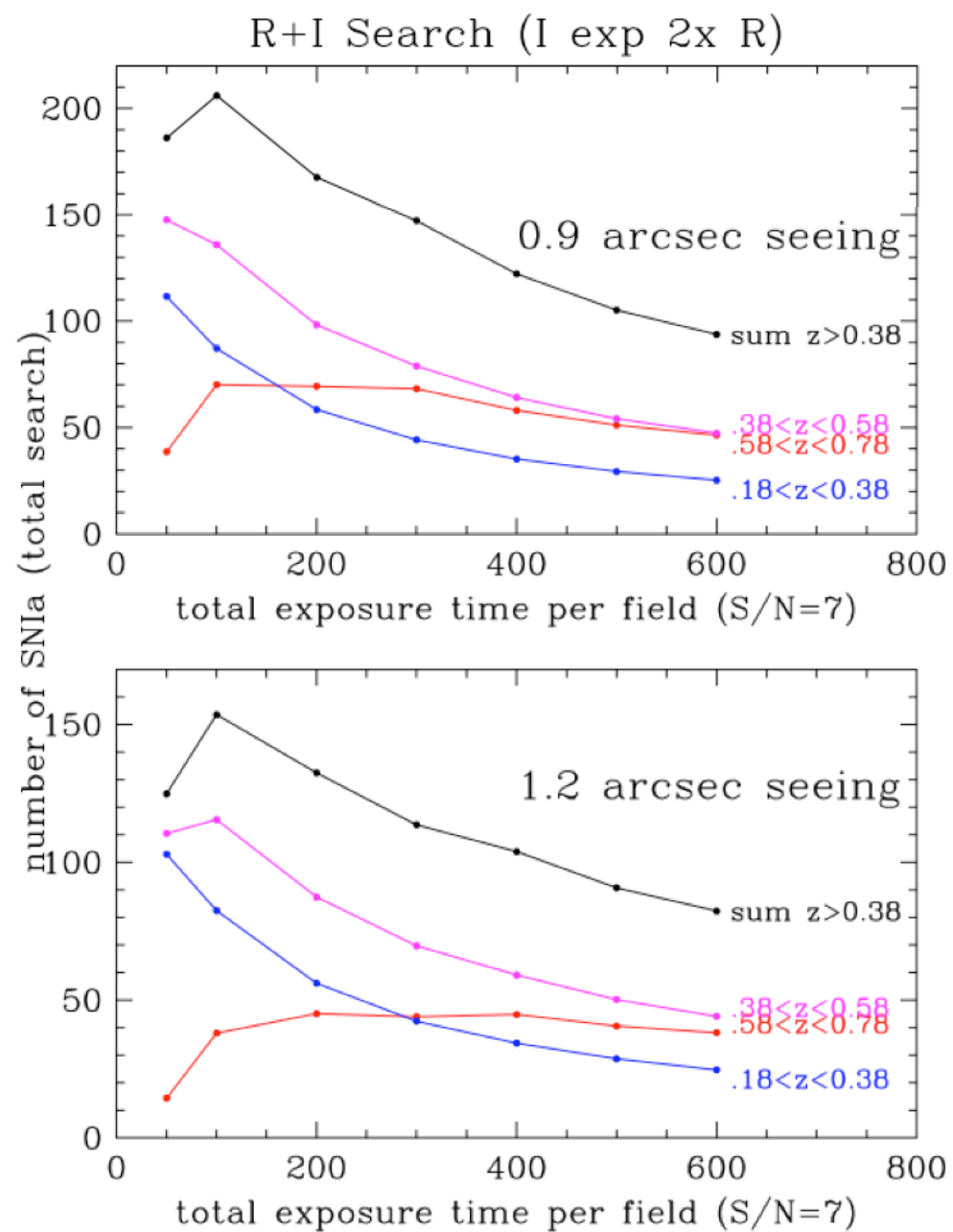
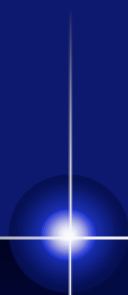
# ESSENCE

- GOAL: Constrain value of  $w$  to within about 10%
- Need  $\sim > 200$  Type Ia SNe: Populate bins of  $\Delta z = 0.1$  in range of  $0.15 < z < 0.75$
- Multiple bands: VRI, R=200s to get out to  $z \sim 0.8$ 
  - Cover redshift range and SN colors
- $\sim 32$  fields =  $\sim 12 \text{ deg}^2$
- 16 fields in half nights every 4 nights for 3 months
- <http://www.ctio.noao.edu/wproject> or <http://www.ctio.noao.edu/essence>

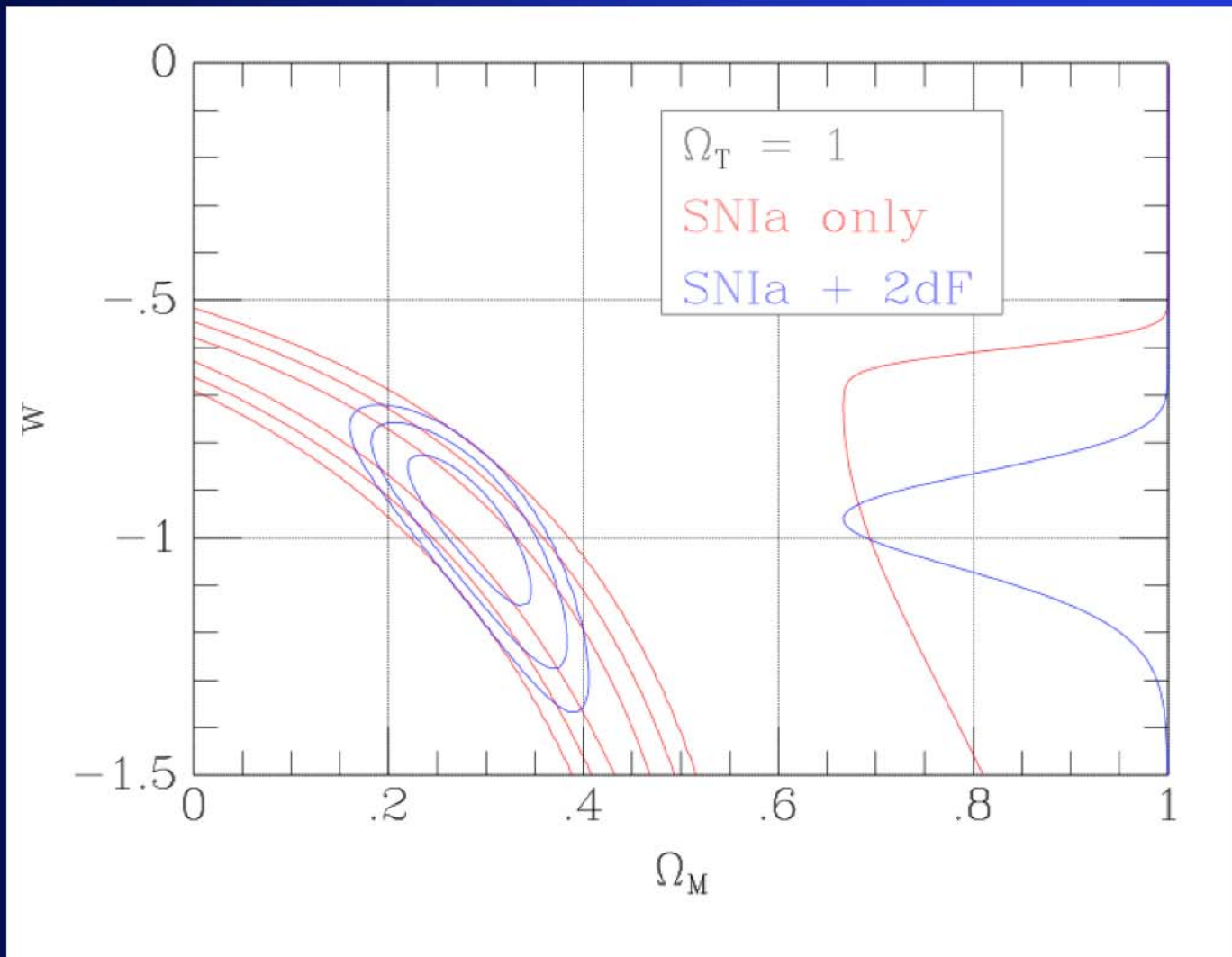
# SN magnitudes as $fn(z)$

TABLE 7  
PEAK MAGNITUDES OF NORMAL SNe Ia

$z$	$B$	$V$	$R$	$I$	$Z$	$J$
0.05 .....	17.37	17.41	17.36	17.82	17.72	(17.70)
0.10 .....	18.94	18.88	18.80	19.20	19.24	(19.30)
0.20 .....	(20.70)	20.29	20.27	20.36	20.75	(20.96)
0.30 .....	(21.94)	21.08	21.07	20.97	21.28	(21.63)
0.40 .....	(23.07)	21.89	21.62	21.55	21.67	21.97
0.50 .....	(24.06)	(22.59)	22.10	22.01	21.96	22.35
0.60 .....	(24.90)	(23.26)	22.59	22.41	22.37	22.79
0.70 .....	(25.53)	(23.97)	23.06	22.70	22.70	23.06
0.80 .....	(26.00)	(24.75)	(23.54)	22.91	23.01	23.11
0.90 .....	(26.39)	(25.46)	(24.06)	23.12	23.17	23.20
1.00 .....	(26.75)	(26.03)	(24.61)	23.48	23.39	23.25
1.10 .....	(27.08)	(26.45)	(25.14)	23.78	23.50	23.38
1.20 .....	(27.39)	(26.78)	(25.66)	(24.19)	23.72	23.59
1.30 .....	(27.67)	(27.07)	(26.13)	(24.59)	24.11	23.72
1.40 .....	(27.95)	(27.35)	(26.55)	(24.91)	(24.39)	23.94
1.50 .....	(28.20)	(27.60)	(26.93)	(25.36)	(24.70)	24.08
1.60 .....	(28.45)	(27.85)	(27.24)	(25.85)	(25.04)	24.18
1.70 .....	(28.68)	(28.08)	(27.51)	(26.30)	(25.32)	24.32
1.80 .....	(28.90)	(28.30)	(27.75)	(26.77)	(25.72)	24.38



# ESSENCE Simulation





# Dark Energy Camera SNe

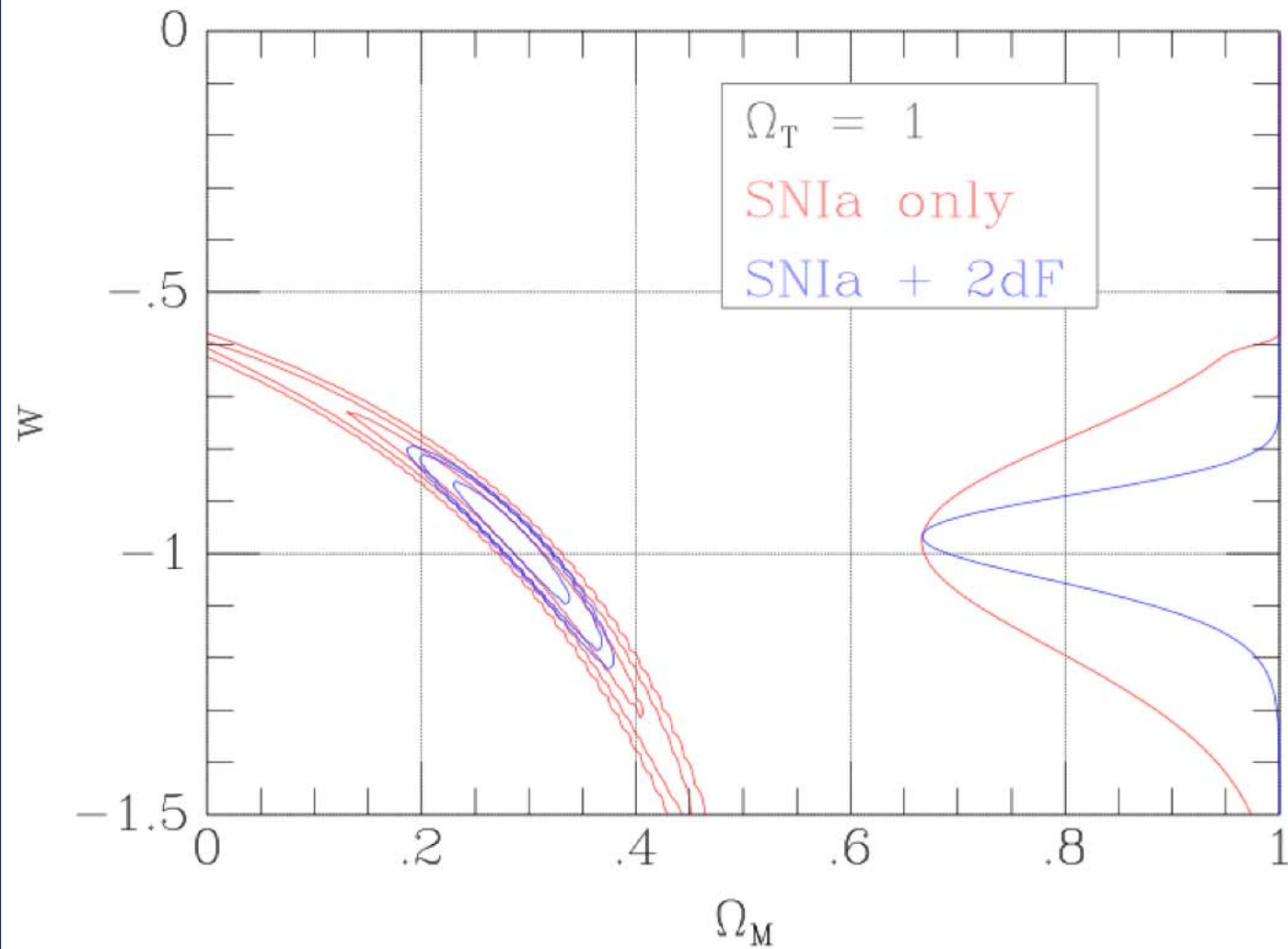
- Science
  - Constrain  $w$  (following simulations)
  - Constrain  $w'$  (idea... no details yet)
- Simulations
  - Use “Tonry toolkit” to simulate a variety of strategies
  - Toolkit takes parameters survey parameters (sampling, depth, area, etc.), generates SN light curves with appropriate photometric errors, fits them, and derives  $w$  contours
  - Gajus Miknaitis ran these simulations for the Dark Energy 2-degree<sup>2</sup> camera

# Simulation “2degA”

- Simulation
  - Used same exposure times as current ESSENCE strategy (R,I=200,400s).
  - Tweaked parameters so five 2 sq. deg. fields are imaged every 3 days for 5 years. Assuming a more sane read time of 20 secs, this requires ~1 hour of imaging every 3 nights.
- Results
  - 1194 SNe, out to  $z \sim 0.8$ .
  - Error in  $w$  looks to be dominated by uncertainty in  $\omega_m$ , taken to be  $\omega_m = 0.278 \pm 0.042$  (i.e. 15%) from 2df



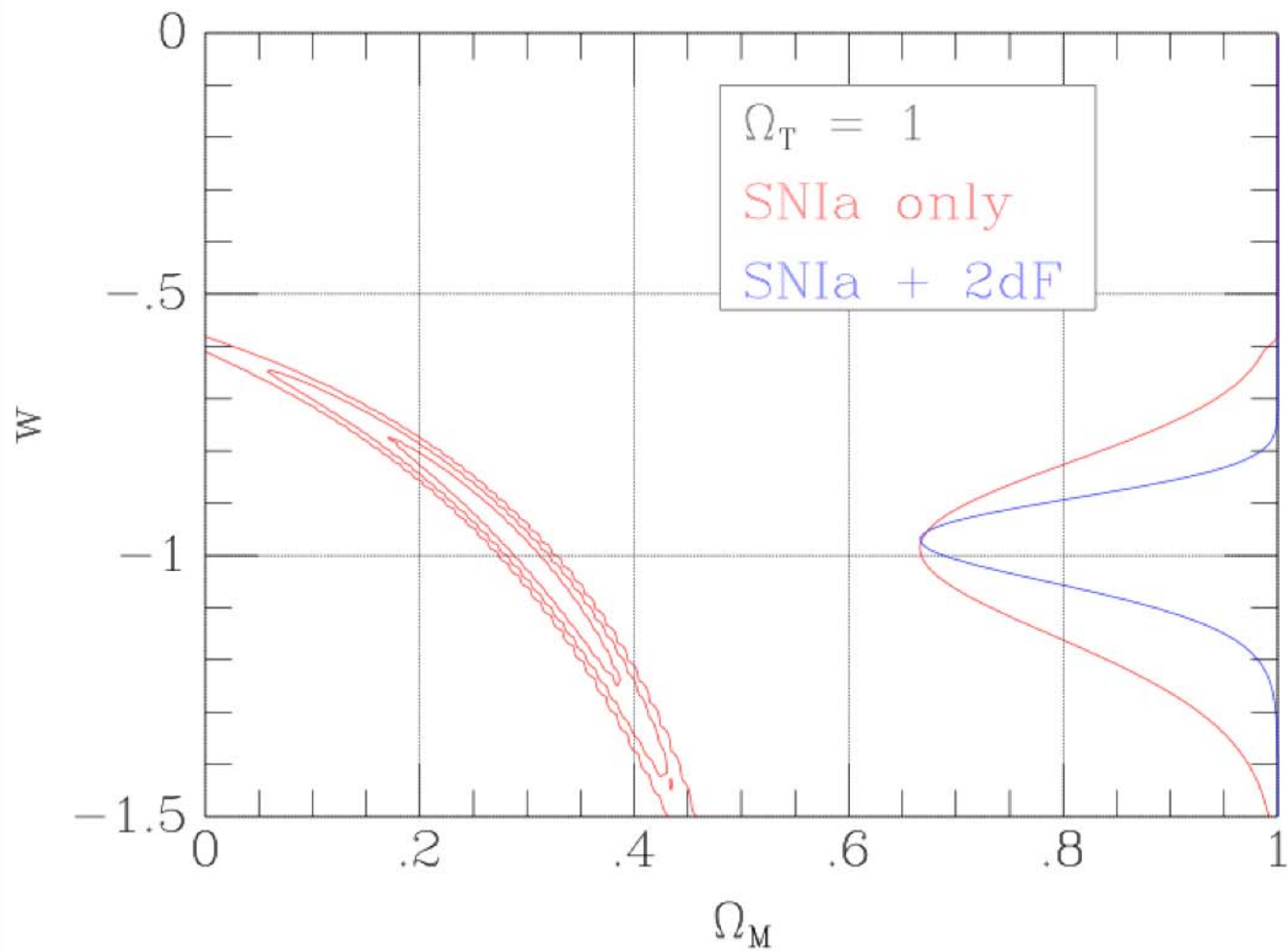
# 2degA results



# Simulation 2degC

- Simulation
  - Same as 2degA (ESSENCE parameters), except double the exposure times in R and I
    - (R,I = 400,800s).
  - This would require ~2 hrs every night.
- Results
  - 1487 SNe, out to  $z \sim 0.85$
  - Very skinny contours, "SN only" constraint approaches joint constraint.

# 2degC Results



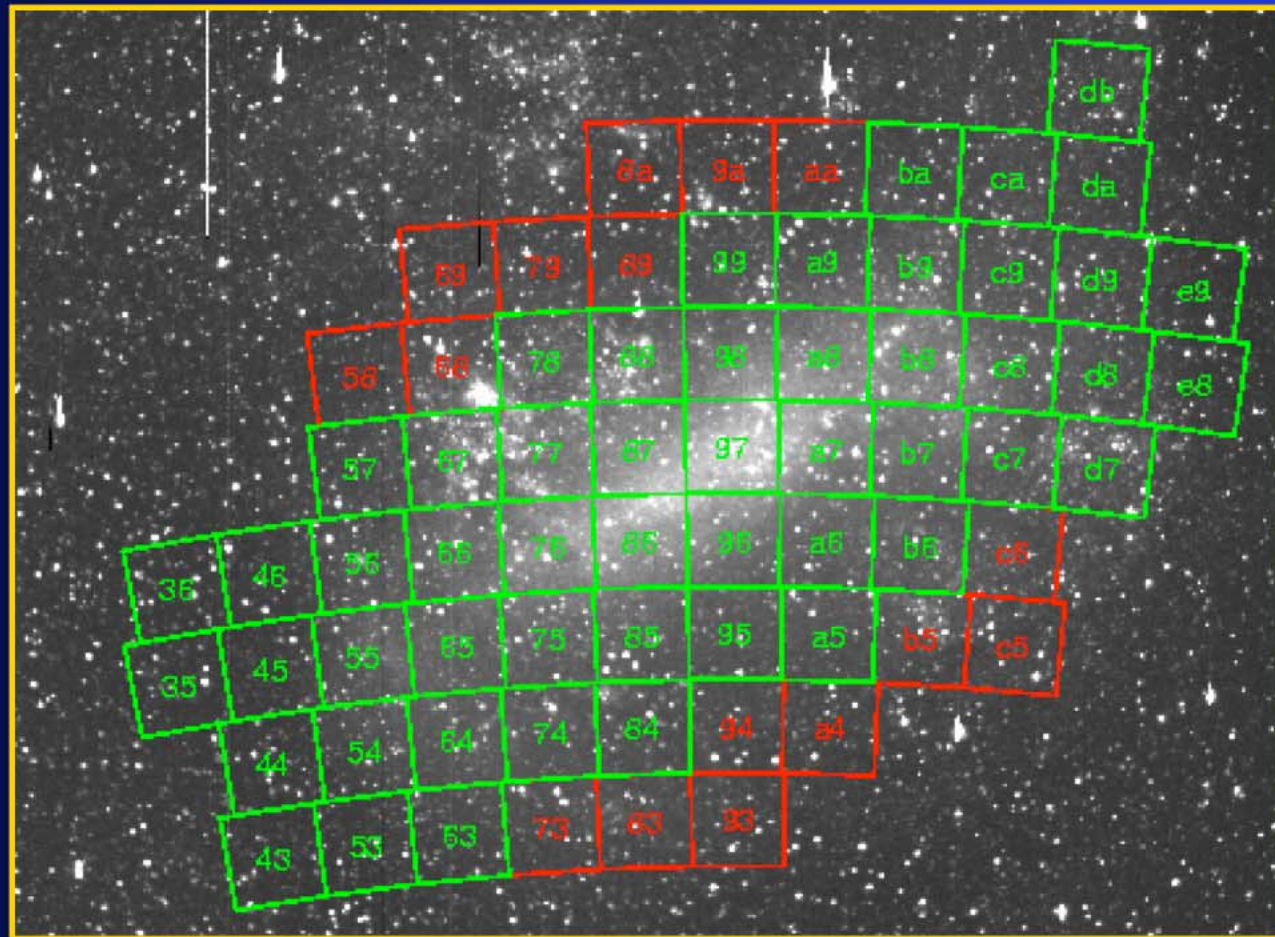
# D.E.C. SN Options

- Include SN strategy in base observations
  - Break multiple exposures in SN-friendly way
- Include deeper “pencil beam” fields within 5000 sq.deg. area
- SN program outside of main survey, but done in conjunction with survey allowing for more flexible scheduling
  - Partial nights
  - Portion of night drifting through nights



# SUPERMACHO

- Study microlensing towards the LMC



# SUPERMACHO

- Goal: Differentiate between whether lenses are in the LMC or Halo
- Need ~50 well-characterized events
- Single band: “VR” = 5200–7200Å
- ~60 fields = ~21 deg<sup>2</sup>
- Half-nights every 2 nights for 3 months (Oct-Dec)
- Exposure times optimized to maximize # of stars, ranging from 20s to 200s
- $\sigma = 0.1$  mag at 23rd
- <http://www.ctio.noao.edu/~supermacho>